

# MULTICOLOR PHOTOMETRY OF THE GALAXIES IN A2255 B THE BEIJING-ARI ONA-TAIWAN-CONNECTICUT SURVEY AND SLOAN DIGITAL SKY SURVEY

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*Received 2002 December 24; accepted 2003 June 17*

## ABSTRACT

We present a multicolor photometric study of the galaxies in the A2255 B field. The data are derived from the Beijing-Ari Ona-Taiwan-Connecticut Survey (BATC) and the Sloan Digital Sky Survey (SDSS). The BATC data were obtained with a CCD camera on the Beijing-Ari Ona-Taiwan-Connecticut Survey telescope, with a field of view of  $58' \times 58'$  and a seeing of  $\sim 0.6''$ . The SDSS data were obtained from the Sloan Digital Sky Survey database. We present the color-magnitude diagrams (CMDs) and color-color diagrams (CCDs) for the galaxies in the A2255 B field. The CMDs show a clear red sequence of galaxies, with a color index of  $r' - i' \sim 2.2$ . The CCDs show a clear separation between the red sequence galaxies and the blue galaxies. We also present the spectral energy distributions (SEDs) for the galaxies in the A2255 B field. The SEDs show a clear peak in the near-infrared region, with a color index of  $r' - i' \sim 2.2$ . The results suggest that the galaxies in the A2255 B field are predominantly red sequence galaxies.

*Subject headings:* galaxies: clusters: rich; galaxies: photometry; galaxies: colors and polarizations; galaxies: distances and redshifts; galaxies: evolution

*On-line material:* color-magnitude diagrams; color-color diagrams; spectral energy distributions

## 1. INTRODUCTION

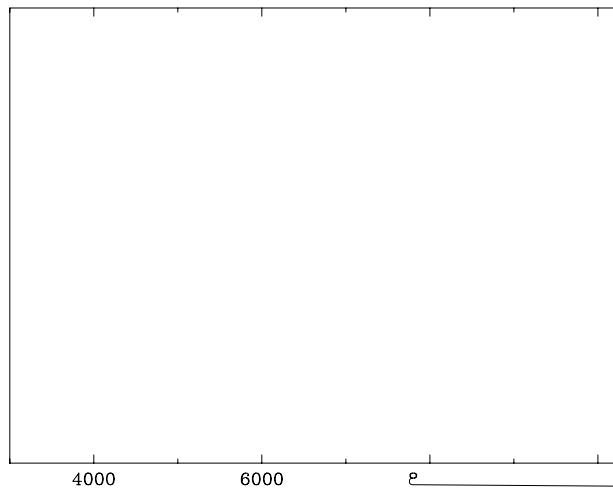
The A2255 galaxy cluster is one of the most prominent clusters in the local universe. It was first discovered by Zwicky (1933) and later identified by de Vaucouleurs & Corwin (1977). The cluster is located at a redshift of  $z \sim 0.068$  and contains a rich population of galaxies. The galaxies in the A2255 cluster are predominantly red sequence galaxies, with a color index of  $r' - i' \sim 2.2$ . The cluster is also rich in blue galaxies, with a color index of  $r' - i' \sim 1.8$ . The blue galaxies are thought to be star-forming galaxies or galaxies in the early stages of their evolution. The A2255 cluster is an excellent laboratory for studying the evolution of galaxies in clusters.

The A2255 cluster has been studied extensively in the past. Zwicky (1933) was the first to identify the cluster. He reported a total mass of  $1.2 \times 10^{15} h^{-1} M_{\odot}$  and a concentration parameter of  $c = 1221$ . Later, de Vaucouleurs & Corwin (1977) studied the cluster in detail. They reported a total mass of  $1.5 \times 10^{15} h^{-1} M_{\odot}$  and a concentration parameter of  $c = 1221$ . The cluster has also been studied by other authors, including Abell (1958), Zwicky & Corwin (1974), and Zwicky & Corwin (1975). The A2255 cluster is a rich source of galaxies for studying the evolution of galaxies in clusters.

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$V \sim 17.5$ ,  $B$  (L 1997),  $V \sim 20.5$ ,  
 $M_V \sim -18.0$ , SDSS,  $F$   
 $(18.0 < m_s < 21.5)$   
 A2255 (BATC)  $B_s$  -  
 -A -T -C (BATC)  $B_s$  -  
 .T (SED) -  
 0.5 (X .2002),  
 60/90 S T S 15 -  
 O (BAO) S -  
 SED 1 2  
 (~1 2)  
 2001). S  
 ) (X  
 2002), SDSS 1996) BATC  
 (F SED SED  
 ) S  
 . B  
 A2255 T  
 BATC  
 BATC SDSS  
 § 3. T SDSS SED  
 SED A2255 § 4, § 5. T  
 § 6, § 7. T  
 $H_0 = 50$   $M^{-1}$   $q_0 = 0.5$



$\sim 3000$  10000 A. T  $B A O S$   
 T CCD (F 1996; 2000; SDSS  
 .2001). F BATC SDSS  
 $u', g', r', i', z'$   
 3560, 4680, 6180, 7500,  
 8870 A. T F 1. BATC SDSS  
 13 BATC A2255. 40  
 . A 36, 113  
 CCD BATC 1", 1. T  
 BATC T S B  
 G S C (PSF) (GSC). T  
 2003). A PSF P 2" (A2255,  
 . F 4 (6", 8"),  
 . A  
 ) (SED)  
 . S  
 T T 1,  
 T relative SED T

## 2. OBSERVATIONS AND DATA REDUCTION

A2255 ( $z \sim 0.0806$ , & R 1999), II III  
 38.3 S (1958),  
 B & M (1970). T BATC  
 T A2255 60/90 /3 S  
 BAO, X  
 900 . A F 2048 x 2048 CCD  
 $\sim 1.0$  T 1"7  
 59 x 59 2,  
 17 08 03.7 17 16 58.6,  
 63°36'31".5 64°35'28".8 (J2000.0). T BATC  
 a-k, m-p, 15

TABLE 1  
THE DETAILS OF THE BATC FILTERS AND OUR OBSERVATIONS

N	F N	$\lambda$ (A)	FWHM (A)	$E^z$ (%)	N I	S (%)	O D	L (%)
1.....		3907	291	10800	9	4.74	5678	20.5
2.....		4540	332	13200	11	5.45	6816	20.5
3.....		4925	374	12000	10	4.32	7227	20.0
4.....		5267	344	11400	10	4.98	7253	20.0
5.....		5790	289	7200	6	3.96	7337	20.0
6.....		6074	308	6950	7	4.18	7319	20.0
7.....		6656	491	12900	12	4.61	6884	19.5
8.....		7057	238	4800	4	4.36	7437	20.0
9.....		7546	192	4800	4	4.62	7386	19.0
10.....		8023	255	14400	12	4.11	7368	19.0
11.....		8484	167	9600	8	5.13	7047	19.0
12.....		9182	247	12000	10	3.81	7355	18.5
13.....		9739	275	12000	10	4.21	6796	18.5

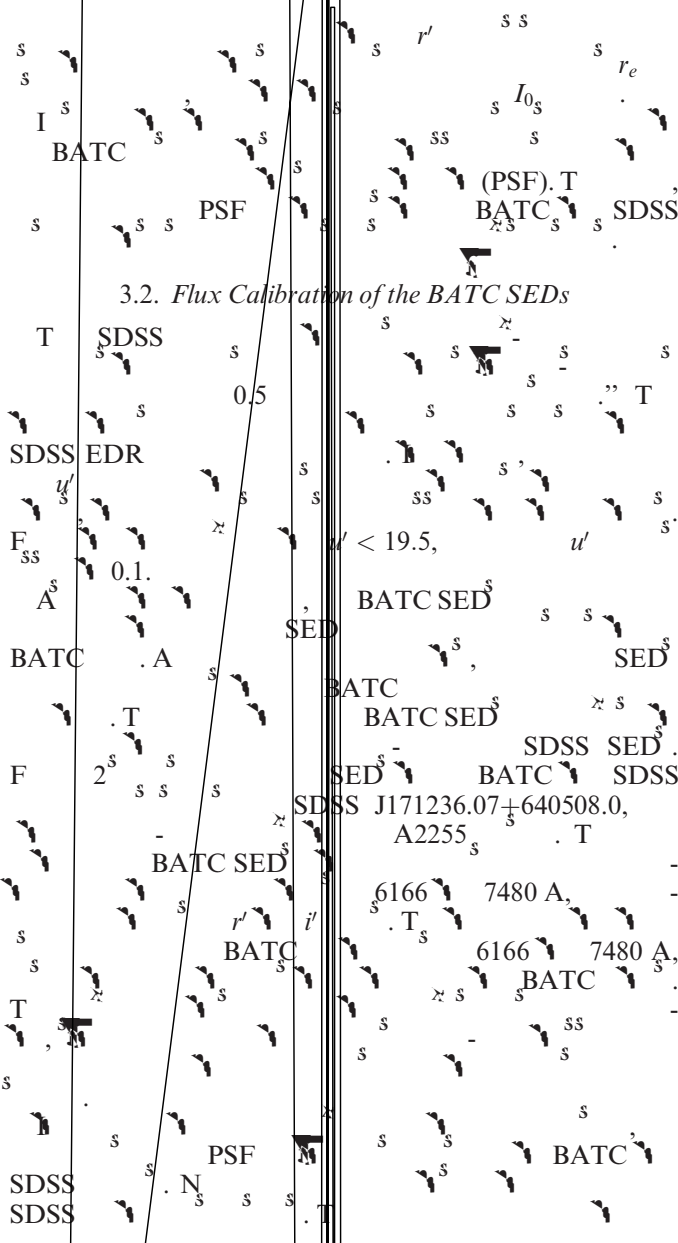
1999). N  
I  
BATC  
(  
3651  
(  
IPAC  
14 (4%)  
X-  
NED-  
254  
BATC  
SDSS

2001).  
SDSS (L  
3.1. Aperture Correction of the SDSS Photometries for Galaxies  
F  
SDSS  
model  
(m)  
V  
(1)  
(1948)  $r^{1/4}$   
V  
C  
1979; C  
model  
(m)  
(PSF)  
(2)  
(3)  
(V  
&  
(m)  
(m)  
$$\Delta m = m - m = -2.5 \frac{\int_0^r 2\pi r I(r) dr}{\int_0^\infty 2\pi r I(r) dr}, \quad (1)$$
  
$$I(r) = \begin{cases} I_0 \{-7.67[(r/r_e)^{1/4}]\}, & 0 < r \leq 7r_e, \\ a_0 I_0 (8 - r/r_e), & 7r_e < r < 8r_e. \end{cases} \quad (2)$$
  
$$I(r) = \begin{cases} I_0 \{-1.68 r/r_e\}, & 0 < r \leq 3r_e, \\ b_0 I_0 (4 - r/r_e), & 3r_e < r < 4r_e. \end{cases} \quad (3)$$
  
$$a_0 = 3.8178 \times 10^{-6} I_0$$
  
$$b_0 = 6.4737 \times 10^{-3} I_0$$
  
model  
SDSS  
 $\Delta m$

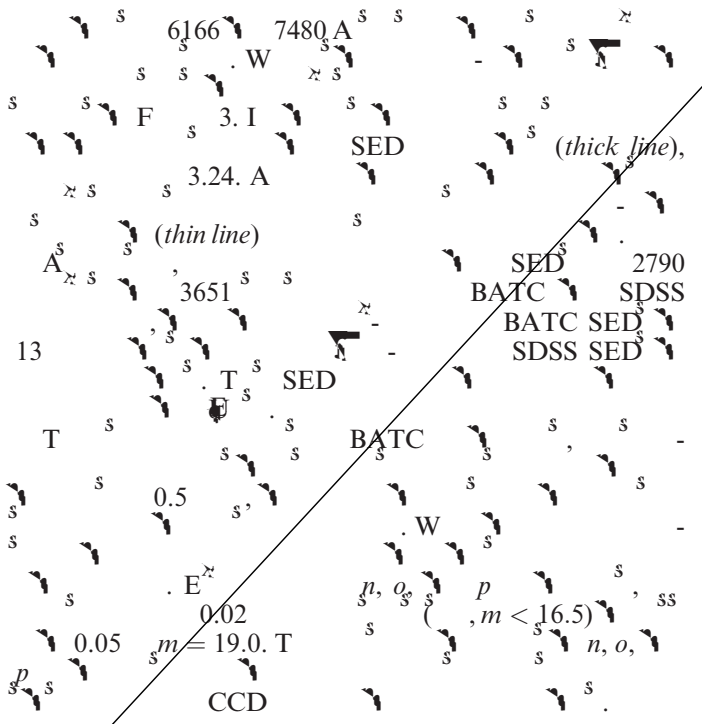
3. METHOD OF COMBINING THE SDSS AND BATC SEDs

I  
SDSS  
BATC SED  
SDSS  
BATC  
 $r = 4 \times 1.7 = 6.8$   
SDSS E D R (EDR;  
2002). T  
SDSS  
A2255  
0.4  
-1  
54

SDSS  
BATC SED  
SDSS  
BATC  
 $r = 4 \times 1.7 = 6.8$   
SDSS E D R (EDR;  
2002). T  
SDSS  
A2255  
0.4  
-1  
54

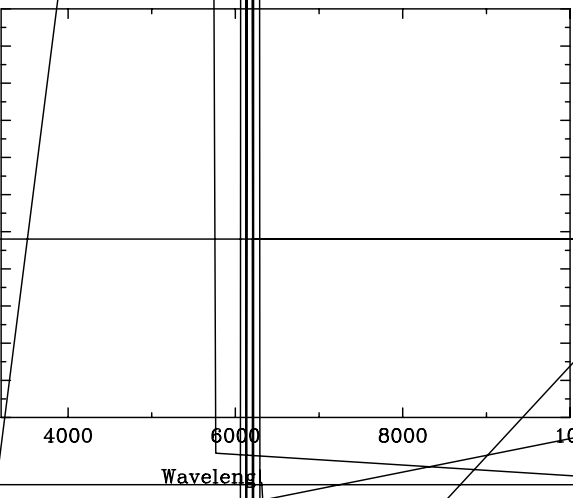


3.2. Flux Calibration of the BATC SEDs



4. ANALYSES OF 254 GALAXIES WITH KNOWN SPECTROSCOPIC REDSHIFTS

4.1. Velocity Distribution and SED Catalog



(1995)<sup>s</sup>, ROSTAT<sup>s</sup>, A2255<sup>s</sup>, (CBI)<sup>s</sup>, (S<sub>BI</sub>)<sup>s</sup>, (B<sub>BI</sub>, F<sub>BI</sub>, & G<sub>BI</sub>)<sup>s</sup>, 1990). B<sub>BI</sub><sup>s</sup>, (1995)<sup>s</sup>, C<sub>BI</sub> = 24330<sup>+203</sup><sub>-265</sub><sup>s-1</sup>, S<sub>BI</sub> = 1240<sup>+203</sup><sub>-129</sub><sup>s-1</sup>, 39<sup>s</sup>, W 254<sup>s</sup>

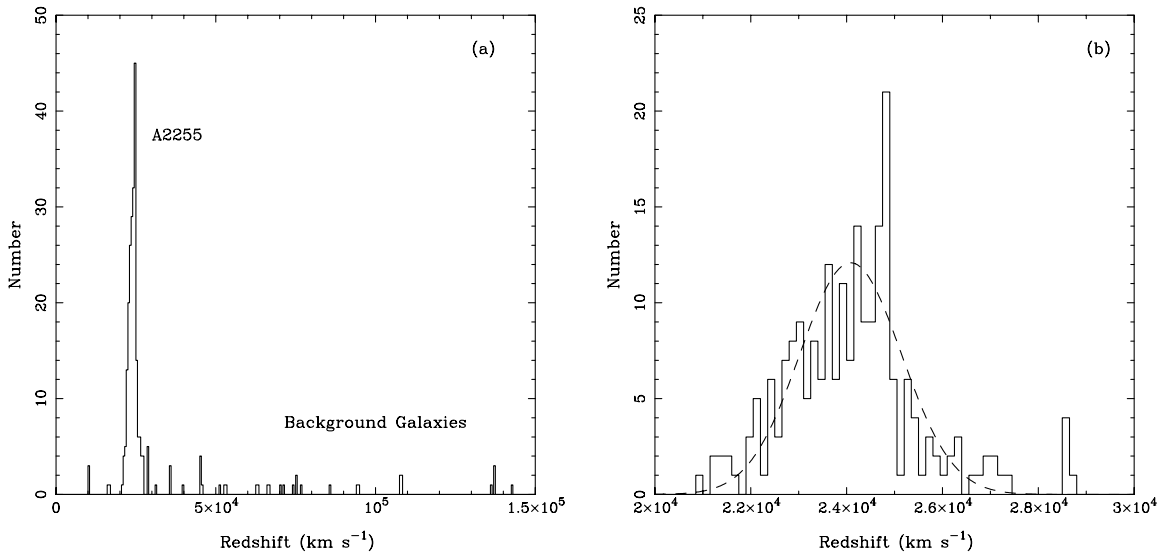
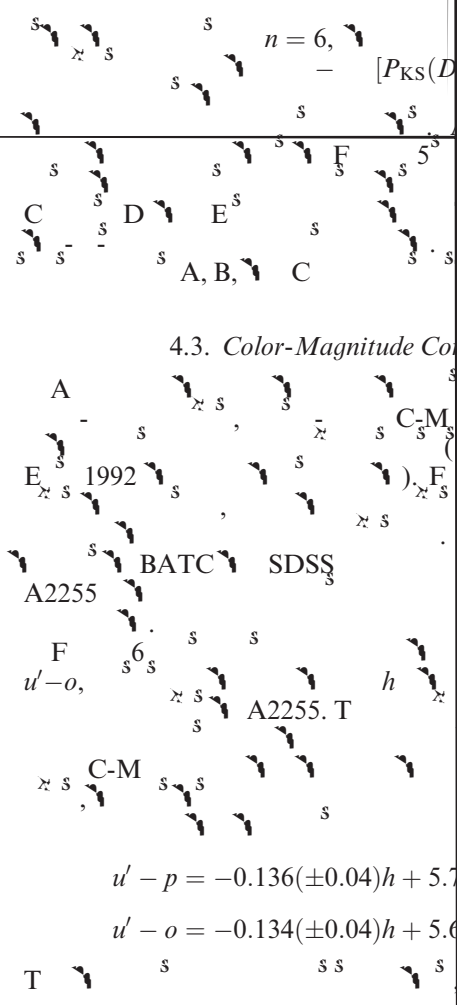
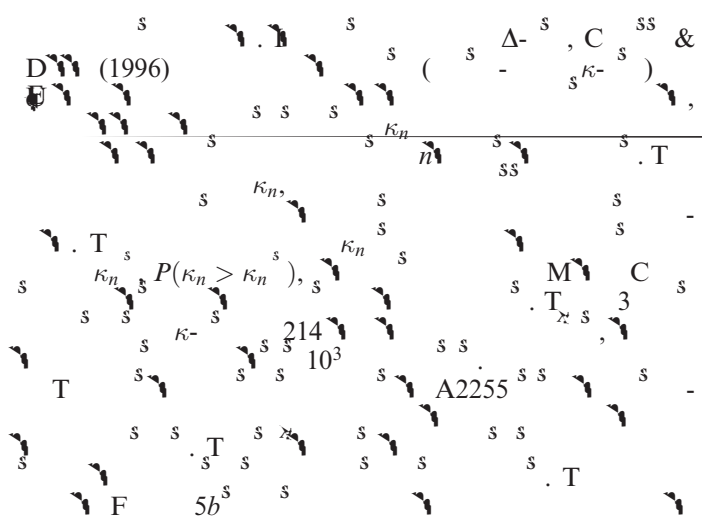
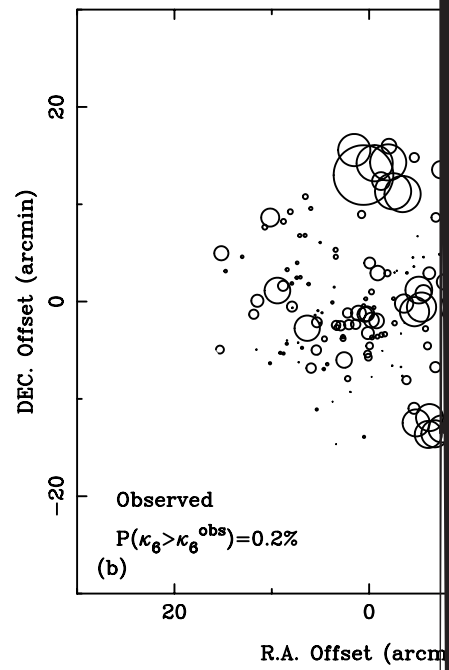
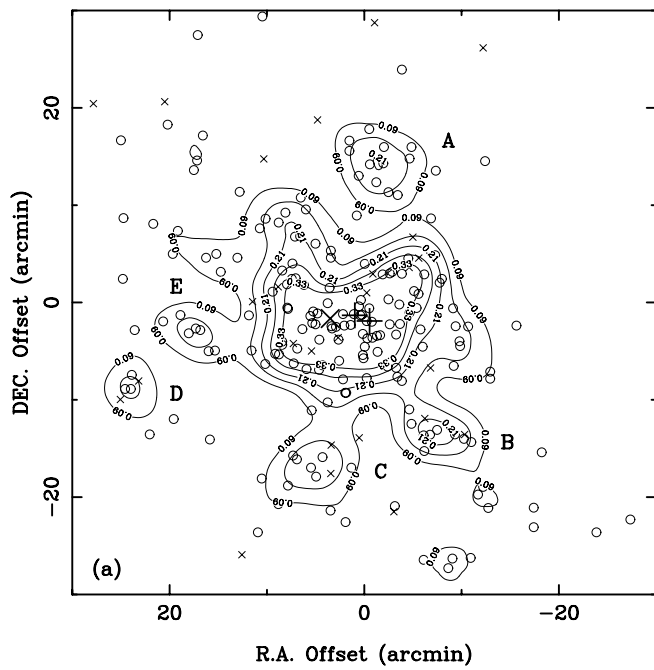


FIG. 4. (a) 254 (b) 214

$214 < cz < 29,000$   
 $C_{BI} = 24025 \pm 89$   
 $S_{BI} = 1315 \pm 86$   
 (1995),  
 $27,454$   
 $94.2\% : 5.8\%$   
 $27,454$   
 $100\%$   
 Column (1).  
 Column (2). R.A. 2000  
 NED.  
 Column (3). D 2000  
 NED.  
 Column (4). S  
 Column (5). P  
 SED.  
 Columns (6)-(18). P  
 BATC . T 0.0

Columns (19)-(23). A  
 SDSS  
 relative SED  
 A2255.  
 4.2. Spatial Distribution and Localized Velocity Structure  
 $214$   
 $168$   
 $46$  (12%)  
 NED-  
 $R.A. = 17 12 31$   
 $= 64^{\circ}05'33''$   
 $23,942$   
 $12$   
 $27,454$   
 $12$   
 $R.A. = 17 12 45$   
 $= 64^{\circ}03'54''$   
 ROSAT  
 $0.15$   
 $0.15$   
 $-2$   
 $-2$   
 (1988)





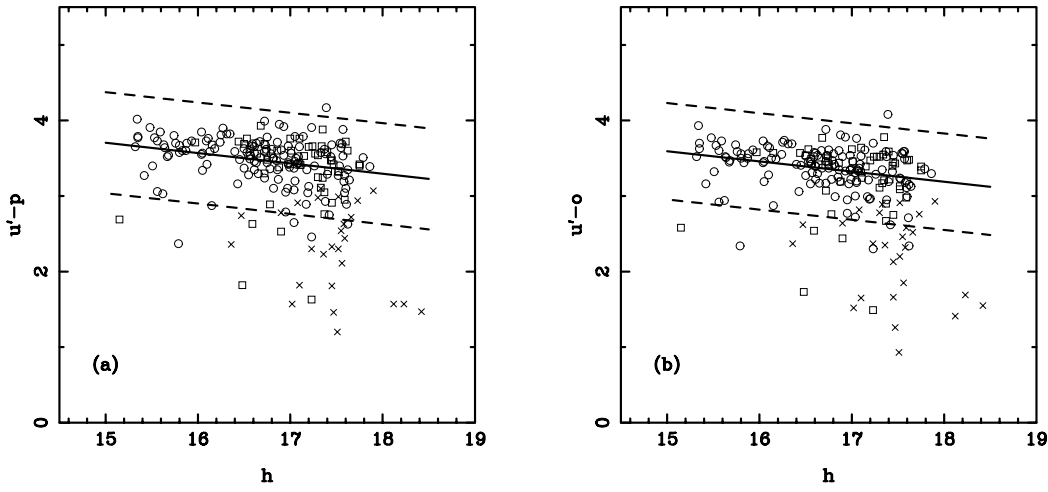


FIG. 6.  $u'-p$  (circles), 51 (open squares), 26 (crosses).  $u'-o$  (circles), 51 (open squares), 26 (crosses).  $h$  (x-axis), 137 (y-axis), 137 (open squares).

(a) (b)  $\sigma_{C.I.} = 0.669$   $0.638$ ,  $95\%$   $M$   $T$   $A2634$   $C-M$   $C-M$  (2001).  $B$   $C$  &  $D$  (1996).

68%  $z \sim 0.075 \pm 0.015$ ,  $z < 0.5$   $A2255$   $A2634$  ( $z \sim 0.03$ )  $BATC$   $SED$   $F$  6  $F$  8  $SED$   $SDSS$   $J171236.07+640508.0$ .  $T$   $SED$

4.4. Application of the Photometric Redshift Technique

$T$   $SED$   $F$   $z$   $SED$   $SED$   $M$   $SED$   $BATC$   $AGN$   $T$   $AGN$   $(13$   $BATC$   $SDSS$   $T$   $254$   $SED$   $SDSS$   $BATC$   $z < 0.5$   $F$   $7$   $(z)$   $254$   $(z)$   $A_V \sim 0.3$   $C$   $(2000)$   $T$   $0.0$   $1.0$   $F$   $7$   $z = z$ ,  $0.01$ .

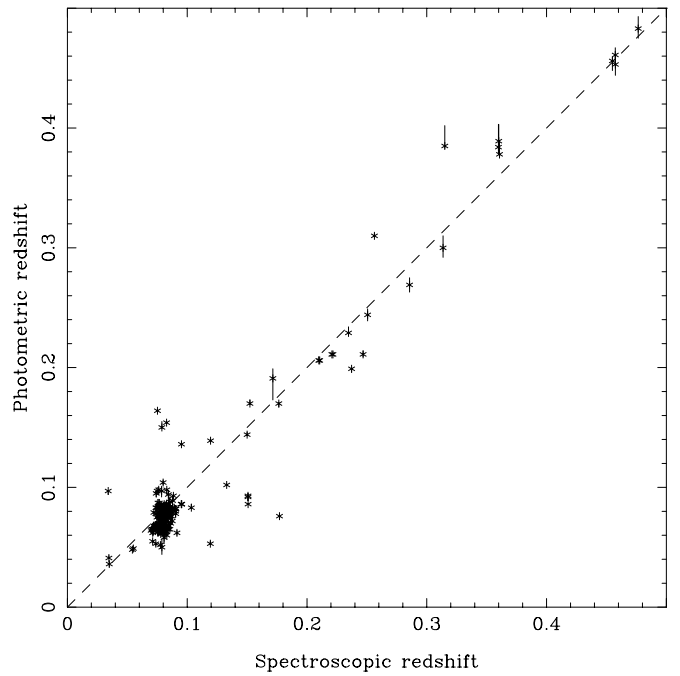


FIG. 7.  $z$  (circles), 254 (crosses),  $z$  (crosses),  $A2255$ .



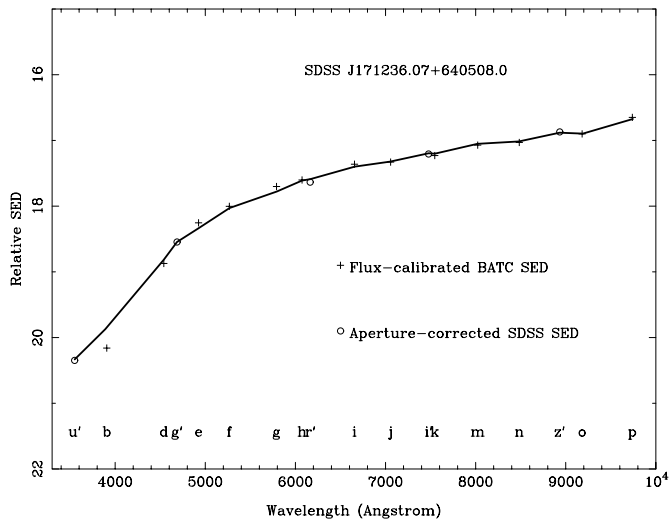
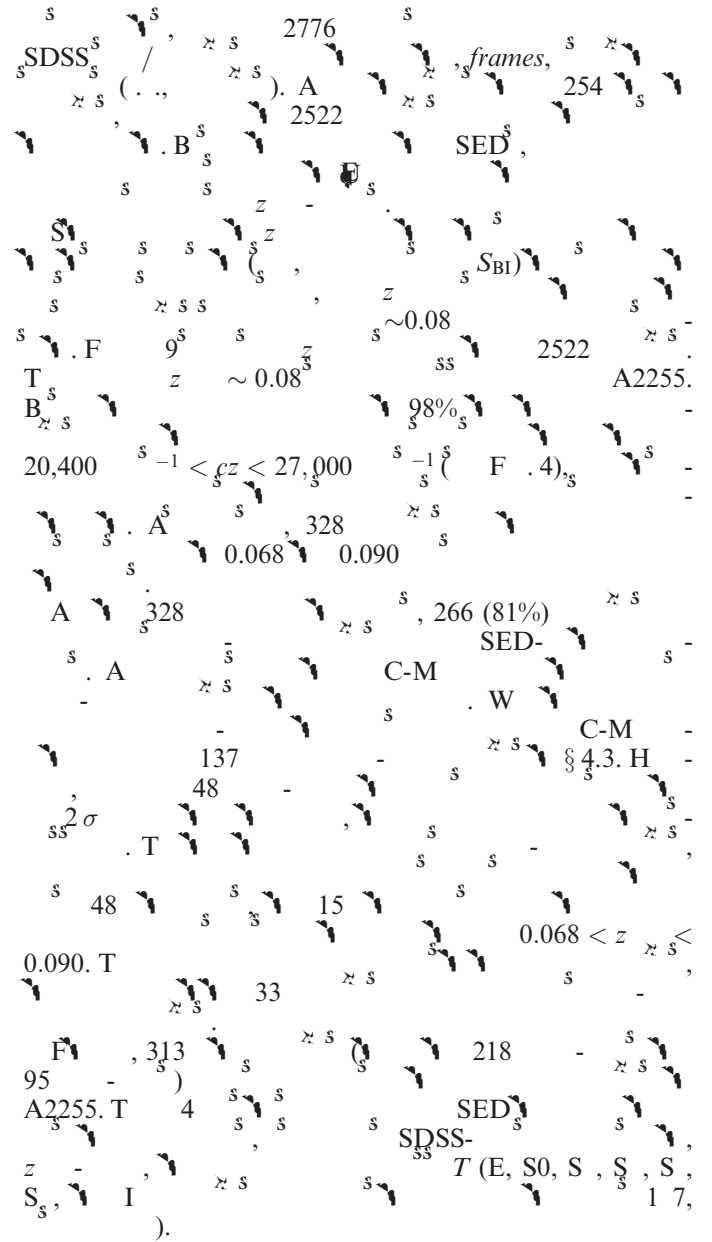


FIG. 8. SED of SDSS J171236.07+640508.0.



5. SED SELECTION OF FAINT CLUSTER GALAXIES

6441 galaxies

BATC

SDSS

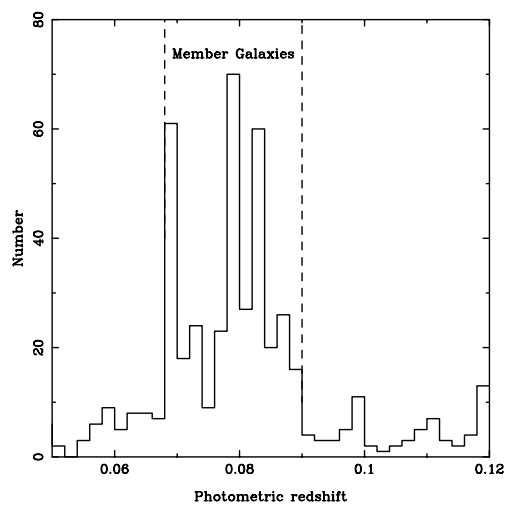
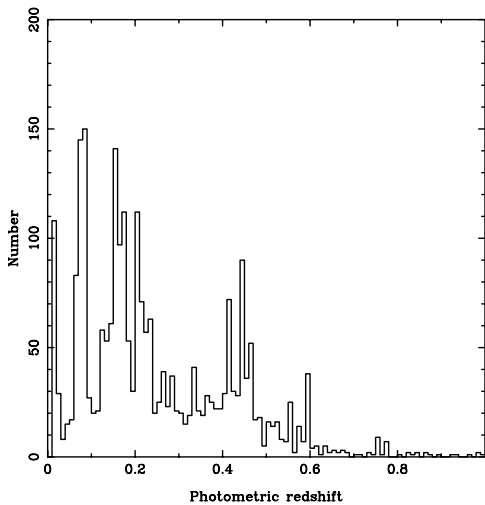


FIG. 9. Photometric redshift distribution.

2522 galaxies

TABLE 4  
THE COMBINED SPECTRAL ENERGY DISTRIBUTIONS OF 313 NEWLY SELECTED MEMBER GALAXIES IN THE REGION OF A2255

N	R.A. (J2000)	D. (J2000)	z	T	b	d	e	f	g	h	i	j	k	m	n	o	p	u'	d'	r'	i'	z'
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)	(12)	(13)	(14)	(15)	(16)	(17)	(18)	(19)	(20)	(21)	(22)	(23)
1.....	17 16 02.36	63 57 55.04	0.079	2	19.85	19.00	18.77	18.62	18.29	18.21	17.98	17.87	17.80	17.58	17.72	17.55	17.41	20.20	18.90	18.16	17.82	17.76
2.....	17 10 16.85	64 21 07.74	0.073	3	20.40	20.01	19.44	19.23	18.91	18.80	18.55	18.46	18.50	18.06	18.21	18.07	18.34	23.02	19.61	18.83	18.43	18.24
3.....	17 12 02.43	63 40 17.14	0.071	1	21.17	20.67	20.53	20.37	19.77	20.03	19.65	19.58	19.32	19.14	19.60	19.12	18.68	22.61	20.64	19.84	19.49	19.31
4.....	17 15 02.65	63 41 03.77	0.086	1	20.32	19.09	18.49	18.26	18.04	17.96	17.71	17.67	17.61	17.47	17.42	17.22	17.02	20.50	18.80	17.96	17.58	17.32
5.....	17 12 45.23	63 41 07.68	0.073	1	22.62	20.62	20.52	20.61	20.25	19.98	19.68	19.70	19.74	19.24	19.32	18.94	19.11	21.99	21.13	20.05	19.61	19.22
6.....	17 10 01.20	63 40 51.23	0.068	1	19.96	19.36	19.15	18.99	18.71	18.76	18.56	18.50	18.50	18.31	18.44	18.19	17.99	20.81	19.41	18.77	18.45	18.26
7.....	17 10 38.30	63 41 00.66	0.087	7	20.55	20.37	20.28	19.92	20.20	20.10	19.83	19.78	19.55	19.56	20.11	20.27	18.94	21.35	20.34	20.04	19.59	19.68
8.....	17 10 21.96	63 41 50.38	0.070	1	23.50	20.84	20.61	20.34	20.07	19.71	19.67	19.35	19.71	19.04	19.27	18.85	19.06	21.94	20.94	19.91	19.45	19.07
9.....	17 14 05.27	63 42 37.08	0.070	3	19.61	23.45	23.06	21.00	20.47	21.02	20.48	20.81	19.77	20.92	19.90	24.62	19.47	21.93	21.41	20.58	20.26	19.87
10.....	17 15 45.57	63 43 12.36	0.088	5	20.48	19.91	19.83	19.92	19.67	19.57	19.52	19.42	19.37	19.45	19.85	19.39	19.83	20.61	19.92	19.50	19.45	19.20

NOTES.  $\emptyset$   
Astrophysical Journal Supplement. A

6. ANALYSES OF THE ENLARGED SAMPLE OF CLUSTER GALAXIES

6.1. Velocity Distributions and Substructures

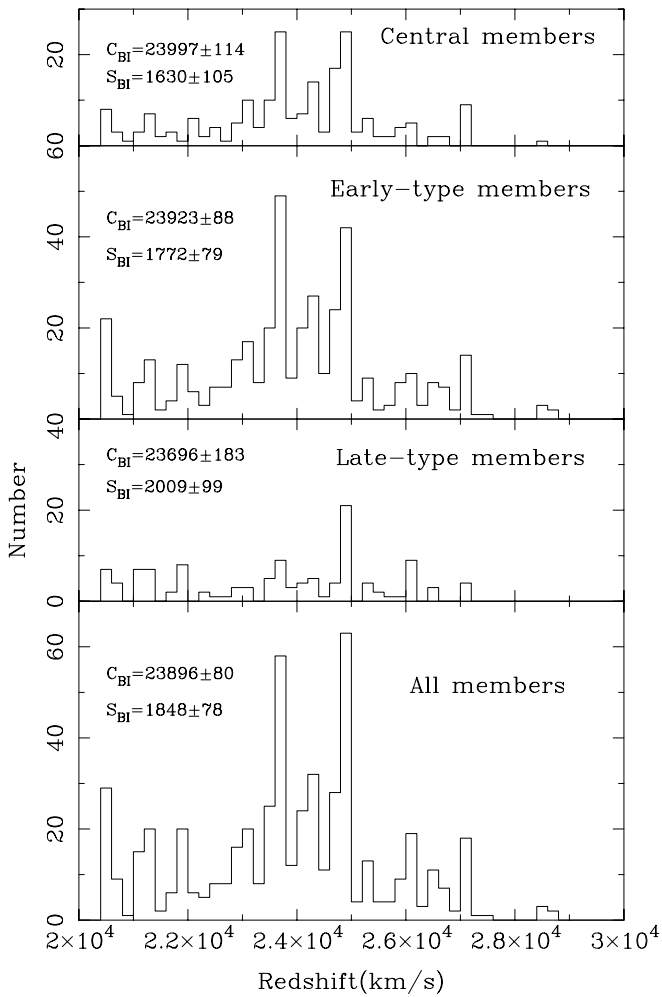
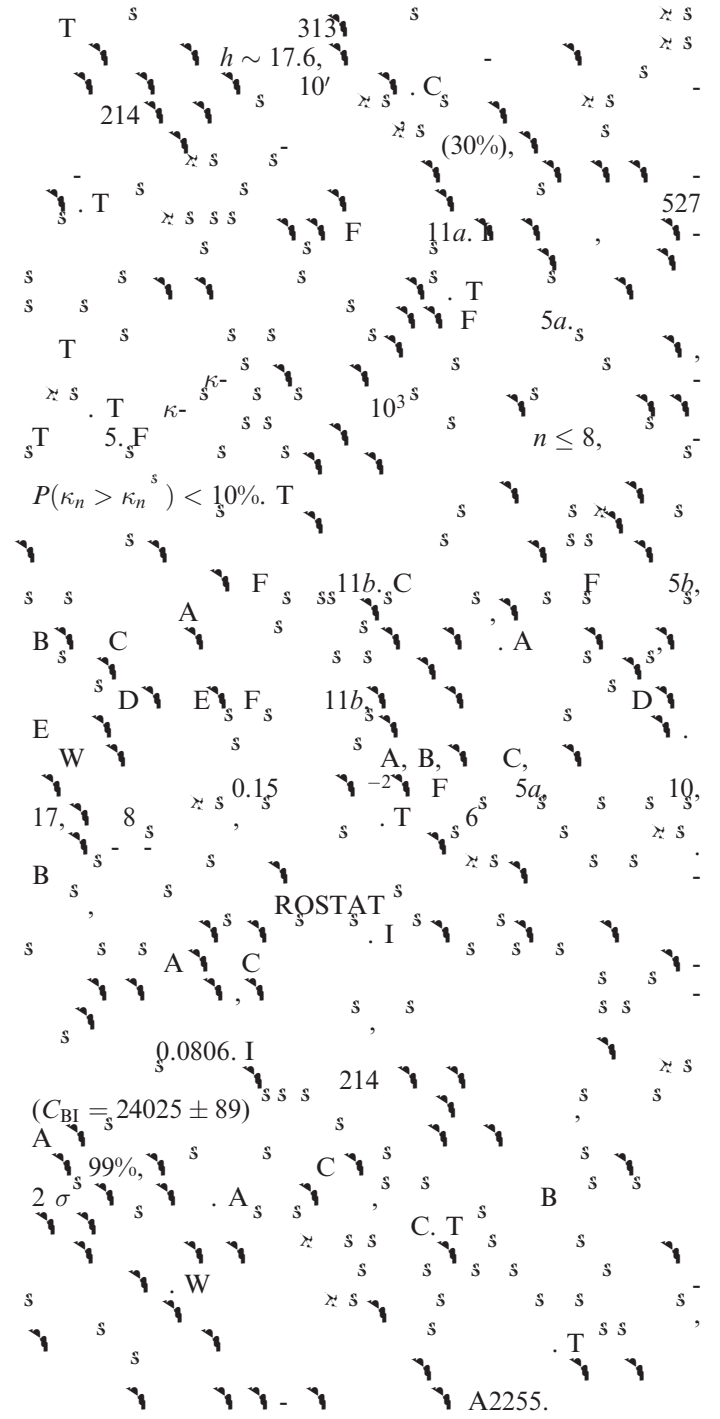
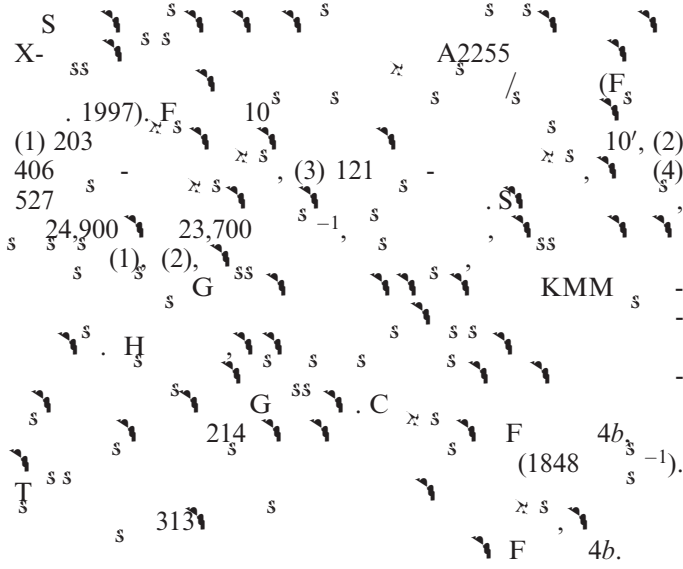
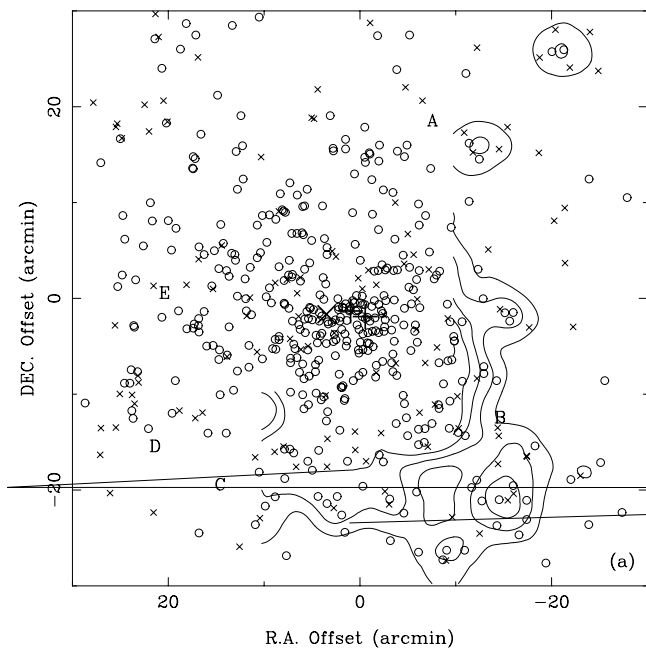


FIG. 10. (1) 203, (2) 406, (3) 121, (4) 527

TABLE 5  
RESULT OF  $\kappa$ -TEST FOR 527 CLUSTER GALAXIES

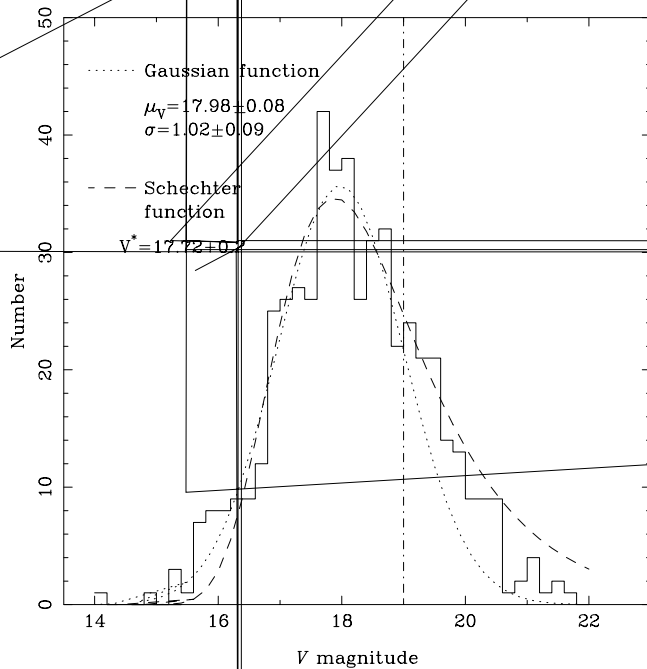
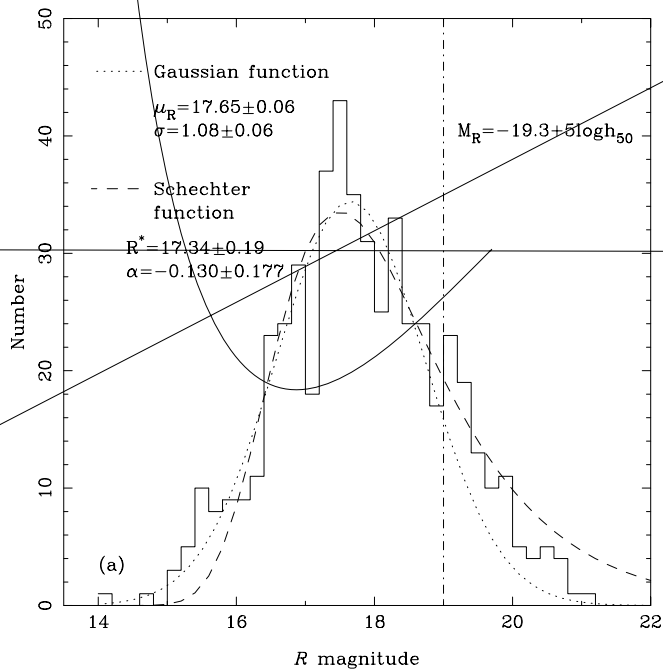
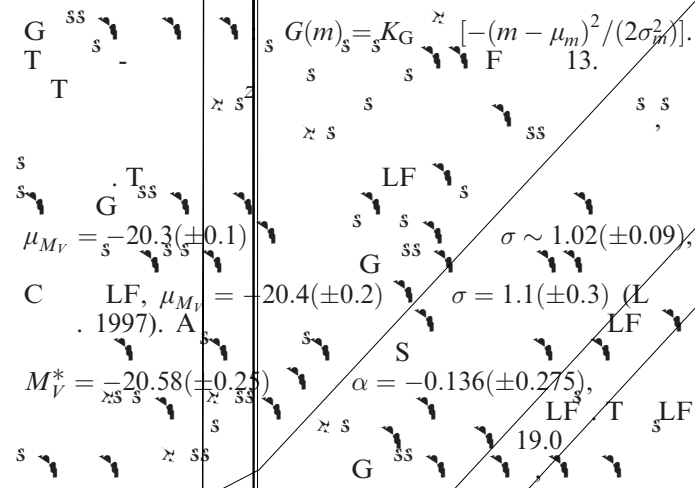
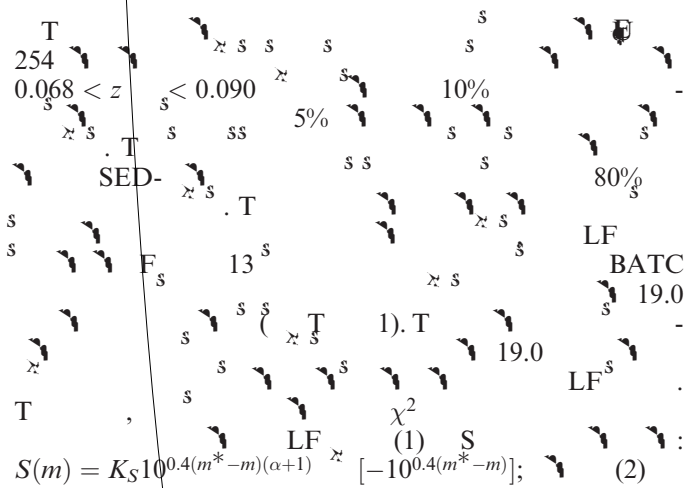
$n$	$P(\kappa_n > \kappa_n^s)$
5.....	0.026
6.....	0.017
7.....	0.057
8.....	0.097
9.....	0.133
12.....	0.174
15.....	0.184
20.....	0.223



6.2. Color Properties and Luminosity Function

T C-M  
 F 12, 20.5, T SDSS  
 15.0  
 2002),  
 4250 Å,  
 C 1992), F  
 T (2001),  
 (LF)  
 R

LF  
 (B 1995; D, A & L 2002). O  
 A2255  
 LF, G  
 1999). F  
 $R = i + 0.1036(\pm 0.055)$   
 $V = g + 0.3292(f - h) + 0.0476(\pm 0.027)$  (2003). F 13  
 A2255. T LF  
 $R \sim 17.5$   $V \sim 17.7$ ,  
 T  
 $(R < 17.5$   $V < 17.8)$  SDSS  
 I  
 LF



A2255.  
 BATC SED  
 BATC

T C-M  
 O V = 19,0, 1  
 527 LF LF

### 7. SUMMAR

T A2255, 60/13  
 90 BATC BAO  
 BATC SED SDSS  
 BATC SDSS model SED  
 254 A2255 SED W  
 90% 0.068 0.090  
 SED W  
 214 SED W  
 A2255 W 137  
 T SED SDSS  
 2522 SED SED- B  
 SED SED- T  
 A 313

S A C  
 I X- SED  
 SDSS BATC SDSS  
 SDSS F

W R. C. B K. M. A W  
 ROSTAT KMM  
 NASA/IPAC E D W  
 J P L (NED), C  
 A T S A S T N  
 F N K B S R S  
 C N S F (NSF)  
 10273007. W P  
 J C J M H W  
 B W P H A

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